

$$= -e^{-u} \left[ \frac{1}{2} s'' + \frac{1}{4} s'^2 + \frac{3}{2} \frac{s'}{r} - \frac{1}{2} \ddot{j} e^u - \frac{3}{4} \dot{j}^2 e^u \right]$$

$$R_3^3 = g^{33} R_{33} = -\frac{e^{-u}}{r^2 \sin^2 \theta} \cdot R_{22} \sin^2 \theta$$

$$= -e^{-u} \left[ \frac{1}{2} s'' + \frac{1}{4} s'^2 + \frac{3}{2} \frac{s'}{r} - \frac{1}{2} \ddot{j} e^u - \frac{3}{4} \dot{j}^2 e^u \right]$$

$$= R_2^2$$

$$R_4^4 = g^{44} R_{44} = \frac{3}{4} \ddot{j} + \frac{3}{4} \dot{j}^2$$

$$\begin{aligned} \therefore R = R_j^j &= R_1^1 + R_2^2 + R_3^3 + R_4^4 \\ &= -2e^{-u} \left( s'' + 2 \frac{s'}{r} + \frac{1}{4} s'^2 \right) + 3(\ddot{j} + \dot{j}^2) \end{aligned}$$

Einstein's field equations are given by

$$R_j^i - \frac{1}{2} g_j^i R + \Lambda g_j^i = -8\pi T_j^i$$

$$\therefore -8\pi T_1^1 = R_1^1 - \frac{1}{2} R + \Lambda$$

$$\Rightarrow 8\pi T_1^1 = -e^{-u} \left( \frac{s'}{r} + \frac{1}{4} s'^2 \right) + \ddot{j} + \frac{3}{4} \dot{j}^2 - \Lambda \rightarrow (7)$$

$$-8\pi T_2^2 = R_2^2 - \frac{1}{2} R + \Lambda$$

$$\Rightarrow 8\pi T_2^2 = -e^{-u} \left( \frac{1}{2} s'' + \frac{1}{2} \frac{s'}{r} \right) + \ddot{j} + \frac{3}{4} \dot{j}^2 - \Lambda \rightarrow (8)$$

$$-8\pi T_3^3 = R_3^3 - \frac{1}{2} R + \Lambda \Rightarrow 8\pi T_3^3 = 8\pi T_2^2$$

$$\Rightarrow 8\pi T_3^3 = -e^{-u} \left( \frac{1}{2} s'' + \frac{1}{2} \frac{s'}{r} \right) + \ddot{j} + \frac{3}{4} \dot{j}^2 - \Lambda \rightarrow (9)$$

$$-8\pi T_4^4 = R_4^4 - \frac{1}{2} R + \Lambda$$

$$\Rightarrow 8\pi T_4^4 = -e^{-u} \left( s'' + 2 \frac{s'}{r} + \frac{1}{4} s'^2 \right) + \frac{3}{4} \dot{j}^2 - \Lambda \rightarrow (10)$$

and  $8\pi T_j^i = 0$  for  $i \neq j$

The assumption of spatial isotropy of 3-space requires that

$$T_1^1 = T_2^2 = T_3^3$$

$\therefore$  from (7) and (8), we have

$$8\pi T_1^1 = 8\pi T_2^2$$

$$\Rightarrow -e^{-u} \left( \frac{1}{2} s'' - \frac{1}{2} \frac{s'}{r} - \frac{1}{4} s'^2 \right) = 0$$

$$\Rightarrow s'' - \frac{s'}{r} - \frac{1}{2} s'^2 = 0$$

$$\Rightarrow \frac{s''}{s'} = \frac{1}{r} + \frac{s'}{2}$$

Integration yields

$$\log s' = \log r + \frac{1}{2} s + \log k_1$$

$$\Rightarrow s' = k_1 r e^{\frac{1}{2} s}$$

$$\Rightarrow \frac{ds}{dr} = k_1 r e^{\frac{1}{2} s} \quad \text{where } k_1 \text{ is constant of integration}$$

$$\Rightarrow e^{-\frac{1}{2} s} ds = k_1 r dr$$

Integrating we get

$$\frac{e^{-\frac{1}{2} s}}{-\frac{1}{2}} = k_1 \frac{r^2}{2} - 2k_2 \quad \text{where } k_2 \text{ is a constant of integration.}$$

$$\Rightarrow e^{-\frac{1}{2} s} = -k_1 \frac{r^2}{4} + k_2$$

$$= k_2 \left( 1 - \frac{k_1}{k_2} \frac{r^2}{4} \right)$$

$$\Rightarrow e^s = \frac{k_2^{1/2}}{\left( 1 - \frac{k_1}{k_2} \frac{r^2}{4} \right)^2}$$

Putting  $-\frac{k_1}{k_2} = \frac{k}{R_0^2} \rightarrow (11)$

since  $R_0^2$  is a constant, which may be positive, negative or infinite, we can express the line-element as

$$ds^2 = dt^2 - \frac{e^{g(x)}}{\left(1 + \frac{k}{4} \frac{r^2}{R_0^2}\right)^2} (dr^2 + r^2 d\theta^2 + r^2 \sin^2 \theta d\phi^2) \quad (12)$$

where  $k = +1, 0, -1$  corresponding to whether the constant  $R_0^2$  given by (11) is positive, infinite or negative.

This completes the derivation of Friedmann - Robertson - Walker (FRW) line-element.

## Dynamical consequences of FRW model:

The Friedmann-Robertson-Walker (FRW) line-element is given by

$$ds^2 = dt^2 - e^{\mu(\pi,t)} (d\pi^2 + \pi^2 d\theta^2 + \pi^2 \sin^2 \theta d\phi^2) \rightarrow (1)$$

with  $\mu(\pi,t) = f(\pi) + g(t)$  and  $e^{f(\pi)} = \frac{1}{\left(1 + \frac{k}{4} \frac{\pi^2}{R_0^2}\right)^2}$

where  $k = +1, 0, -1$  corresponding to whether the constant  $R_0^2$  is positive, infinite or negative.

Assuming the universe to be filled with a highly ideal fluid corresponding to an isotropic and homogeneous distribution of matter having average density  $\rho_0$  and average pressure  $p_0$ , in comoving

coordinate system with  $x^1 = \pi$ ,  $x^2 = \theta$ ,  $x^3 = \phi$ ,  $x^4 = t$ , the energy momentum tensor  $T_j^i$  given by  $T_j^i = (\rho_0 + p_0) u^i u_j - g_j^i p_0$  has the component

$$\left. \begin{aligned} T_1^1 = -p_0, T_2^2 = -p_0, T_3^3 = -p_0 \\ T_4^4 = \rho_0, \text{ and } T_j^i = 0 \text{ for } i \neq j \end{aligned} \right\} \rightarrow (3)$$

Therefore, for the line element (1) the Einstein field equations

$$R_j^i - \frac{1}{2} g_j^i R + \Lambda g_j^i = -8\pi T_j^i \rightarrow (4)$$

gives

$$8\pi(-p_0) = -e^{-\mu} \left( \frac{f'}{\pi} + \frac{1}{4} f'^2 \right) + \ddot{g} + \frac{3}{4} \dot{g}^2 - \Lambda \rightarrow (5)$$

$$8\pi(-p_0) = -e^{-\mu} \left( \frac{1}{2} f'' + \frac{1}{2} \frac{f'}{\pi} \right) + \ddot{g} + \frac{3}{4} \dot{g}^2 - \Lambda \rightarrow (6)$$

$$8\pi \rho_0 = -e^{-\mu} \left( f'' + 2 \frac{f'}{\pi} + \frac{1}{4} f'^2 \right) + \frac{3}{4} \dot{g}^2 - \Lambda \rightarrow (7)$$